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# The zeroth law of black hole thermodynamics revisited

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## Zeroth law of thermodynamics

... defines a notion of equilibrium:



... and of temperature. (intensive quantity!)

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Black hole thermodynamics I

Analogy to ordinary TD:

energy...  $U \leftrightarrow M$  ... mass temperature...  $T \leftrightarrow \kappa$  ... surface gravity entropy...  $S \leftrightarrow A$  ... horizon area .....  $\leftrightarrow$  .....

First law of BH TD:

 $dM = TdS + \Omega dJ + \Phi dQ$ 

Smarr relation: [1]

 $(D-3)M = (D-2)TS + (D-2)\Omega J + (D-3)\Phi Q$ 

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## Black hole thermodynamics II

Conjugate variables: intensive  $\leftrightarrow$  extensive!

$$T = \frac{\partial M}{\partial S}, \quad \Omega = \frac{\partial M}{\partial J}, \quad \Phi = \frac{\partial M}{\partial Q}$$

Fundamental relation:

$$S(M, J, Q) = \pi \left( 2M^2 - Q^2 + 2\sqrt{+M^4 - M^2Q^2 - J^2} \right)$$

 $\implies$  entropy representation:

$$\left\{S, M, J, Q, \frac{1}{T}, -\frac{\Omega}{T}, -\frac{\Phi}{T}\right\}$$

#### Black hole thermodynamics: AdS extension

AdS black holes: [1, 2]

pressure...  $P \leftrightarrow \Lambda$  ... cosmological constant volume...  $V \leftrightarrow V_{th}$  ... thermal volume of the bh

First law: derived from geometric properties.

 $dM = TdS + \Omega dJ + \Phi dQ + VdP \equiv dH$ 

 $\implies$  mass equals enthalpy: M = H = U + PV



Phase transitions in black hole systems?

• Hawking-Page phase transition: [3, 4]

between thermal AdS space and large black hole solution. Calculating the free energy!

• Kerr-AdS: [1, 2]

Van der Waals-type phase transition in P-V-diagram! Swallow tail in Gibbs free energy.



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Questions:		

# Is the P-V-diagram at constant T transferable to black hole thermodynamics?

# Is T a valid intensive quantity to measure thermodynamic equilibrium?

Suppose:

S... thermodyn. potential,  $\{q^i\}$ ... extensive vars.,  $\{p_i\}$ ... intensive vars.

### (Quasi-)Homogeneity in thermodynamics

Quasi-homogeneity: [5, 6] of degree r and type  $\beta$ 

$$S(\lambda^{\beta_1}q^1,...,\lambda^{\beta_n}q^n) = \lambda^r S(q^1,...,q^n)$$

Homogeneity of degree 1:  $S(\lambda U, \lambda V, \lambda N) = \lambda S(U, V, N)$ 

Intensive thermodynamic variables:

$$p_i\left(q^i\right) \equiv rac{\partial S\left(q^i\right)}{\partial q^i} \implies p_i\left(\lambda^{\beta_j}q^i\right) = \lambda^{r-\beta_i}p_i(q^i)$$

... quasi-homogeneous of degree  $r - \beta_i$ .

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### Euler's theorem

For quasi-homogeneous functions holds:

$$rS(q^{j}) = \sum_{i=1}^{n} \beta_{i} \left[ q^{i} p_{i}(q^{j}) \right]$$

Smarr relation for black holes  $\equiv$  Euler relation

$$S = \frac{1}{2} \frac{1}{T} M - \frac{1}{2} \frac{\Omega}{T} J - \frac{\Phi}{T} Q + \frac{1}{2} \frac{P}{T} V$$
  
read off degree and type!

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### **Gibbs-Duhem relation**

... obtained by combining Smarr relation with first law. Usually:

$$Ud\left(\frac{1}{T}\right) - Vd\left(\frac{P}{T}\right) + Nd\left(\frac{\mu}{T}\right) = 0$$

#### Generalized Gibbs-Duhem:

$$\sum_{i=1}^{n} \left[ \underbrace{(\beta_i - r)p_i(q^j)dq^i}_{(1)} + \underbrace{\beta_i q^i dp_i(q^j)}_{(2)} \right] = 0$$

 $\implies$  (1) Variation of extensive quantities  $q^i$  $\implies$  (2) Usual variation of intensive quantities  $p_i$ 

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Mathematical inconsistency

Is the generalized Gibbs-Duhem relation fulfilled in the quasi-homogeneous case?

Definition of equilibrium:

 $dp_i = 0 \Longrightarrow (2) = 0$ 

- Homogeneous TD:  $\beta_i = r = 1 \Longrightarrow (1) = 0$
- Quasi-homogeneous TD:

$$\beta_i, r \neq 1 \Longrightarrow (1) \neq 0$$

(1) related to latent heat in phase transitions!

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### Generalized intensities

# **Redefine equilibrium!**

$$\tilde{p}_i(q^i) \equiv \left[ \left(q^i\right)^{\beta_i - r} \right]^{1/\beta_i} p_i(q^i)$$

... obtained by solving the generalized Gibbs-Duhem relation.

- ... one particular solution, more general intensities possible.
- ... quasi-homogeneous of degree 0.
- ... reduce to  $p_i$  in the case  $\beta_i = r = 1$ .

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Generalized Gibbs-Duhem relation

Rewrite the generalized Gibbs-Duhem relation:

$$\sum_{i=1}^{n} \left[ \left(\beta_{i} - r\right) p_{i}(q^{j}) dq^{i} + \beta_{i} q^{i} dp_{i}(q^{j}) \right] =$$
$$\equiv \sum_{i=1}^{n} \beta_{i} \left(q^{i}\right)^{r/\beta_{i}} \underbrace{d\tilde{p}_{i}(q^{j})}_{=0} = 0$$

Fulfilled by new definition of equilibrium!

The simplest case: not quite significant?

Schwarzschild entropy:  $S(M) = 4\pi M^2 \Rightarrow r = 2, \ \beta = 1$ 

Usual temperature:

$$\frac{1}{T} = \frac{\partial S}{\partial M} = 8\pi M$$

... homogeneous of degree 1! Not truly intensive!



## The popular case: Kerr-AdS

Fundamental potential: H(S, P, J = 1)

- "Intensive" quantities T, P not (quasi-)homogeneous of degree 0.
- Explicit calculation: generalized Gibbs-Duhem not fulfilled
- Phase transition: swallow tail in Gibbs free energy G(T, P, J) = U TS + PV



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The popular case: Kerr-AdS

Introduce new intensive quantities:

$$\tilde{T} = TU$$
 and  $\tilde{P} = PUV^{1/3}$ 

#### ... quasi-homogeneous of degree 0!



Another interesting case: Hawking-Page?

PT in AdS black holes: J = 0

But: *S* and *V* are not independent variables!

$$\left(\frac{3V}{4\pi}\right)^2 = \left(\frac{S}{\pi}\right)^3$$

 $\implies P(V,T)$  and G(T,P) cannot be calculated!

# PT structure has been found in different way! Method not applicable?

## Conclusions

 $\implies$  Introduction of true intensities is important in quasi-homogeneous thermodynamics.

- Revise notion of TD equilibrium in general?
- Thermodynamic response functions?
- Extend methodology?
- Fundamental meaning?

# Thanks!

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#### **References** I

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#### B P Dolan.

Where is the PdV term in the fist law of black hole thermodynamics? arXiv preprint arXiv:1209.1272, 2012.



#### D Kastor, S Ray, and J Traschen.

Enthalpy and the mechanics of AdS black holes. Classical and Quantum Gravity, 26(19):195011, 2009.



#### S. W. Hawking and D. N. Page.

Thermodynamics of black holes in anti-de sitter space. *Commun. Math. Phys.*, 87:577–588, 1983.



#### T. Hartman.

Gr lecture notes. 2015.



F Belgiorno.

Quasi-homogeneous thermodynamics and black holes. Journal of Mathematical Physics, 44(3):1089–1128, 2003.



#### F Belgiorno and SL Cacciatori.

General symmetries: From homogeneous thermodynamics to black holes. *The European Physical Journal Plus*, 126(9):1–19, 2011.